

ACES Payload system test

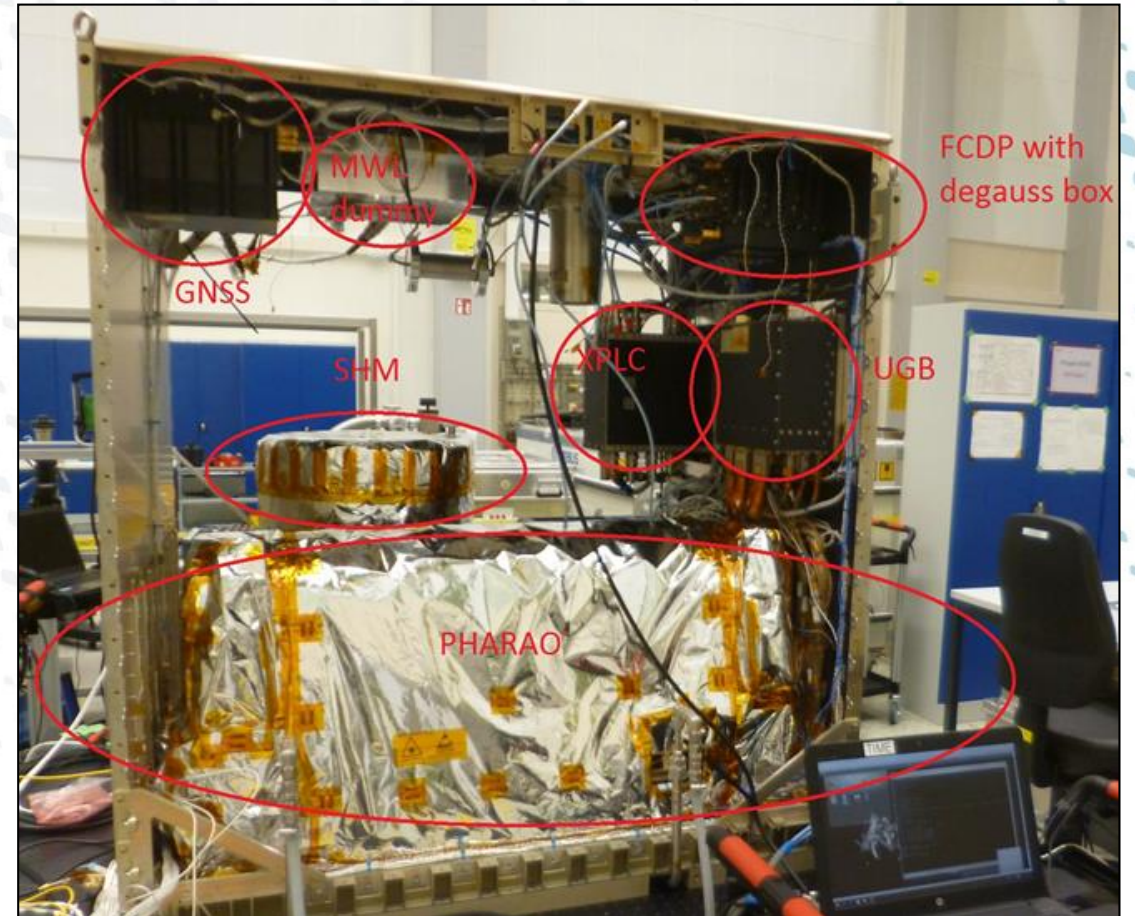
Scientific relevant results

DEFENCE AND SPACE

Silvio Koller Dev. Engineer for fundamental physics,
20.10. 2022

Agenda

- SHM PHARAO FCDP standalone
- Stand alone environmental
- STSL/AMPM
- STSL magnetic environment
- LTSL
- Orbital simulation
- Lessons learned

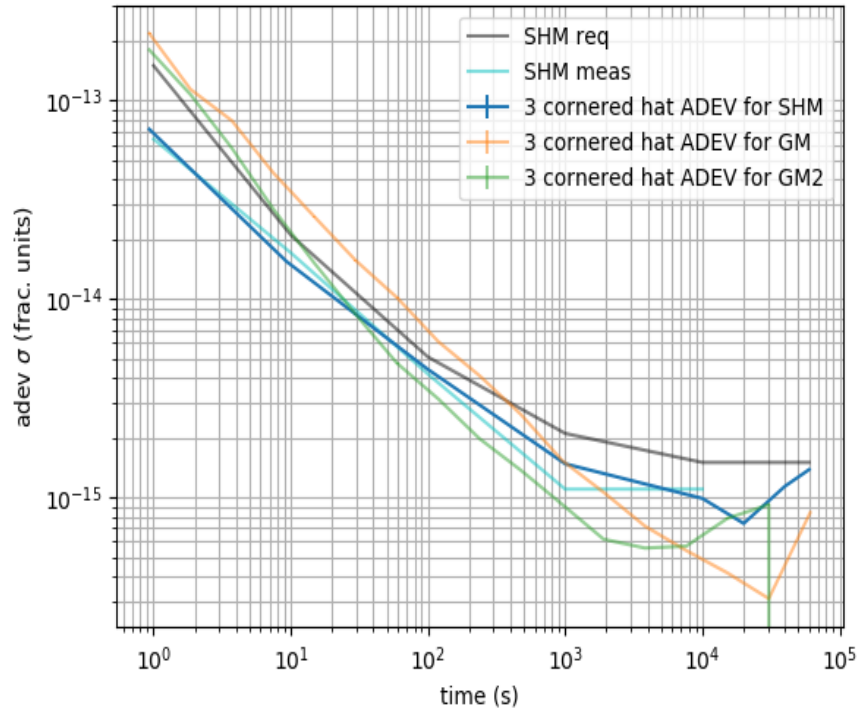


Stand alone measurements:

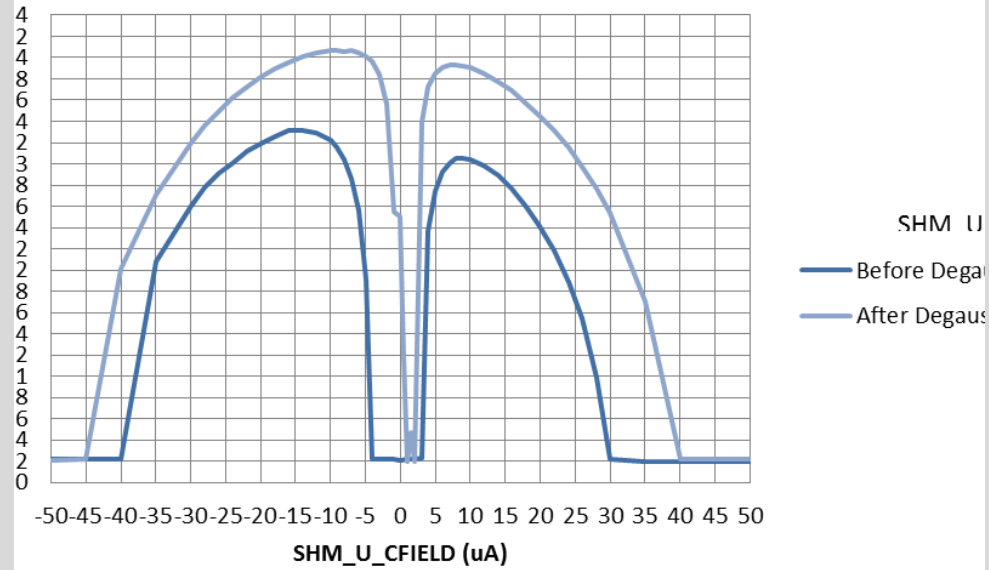
Calibration of active Hydrogen Maser (SHM), Cs clock (PHARAO) FCDP.

Not included is the Microwave link (MWL)
see Achim Helm

three cornered hat
SHM_stand_alone



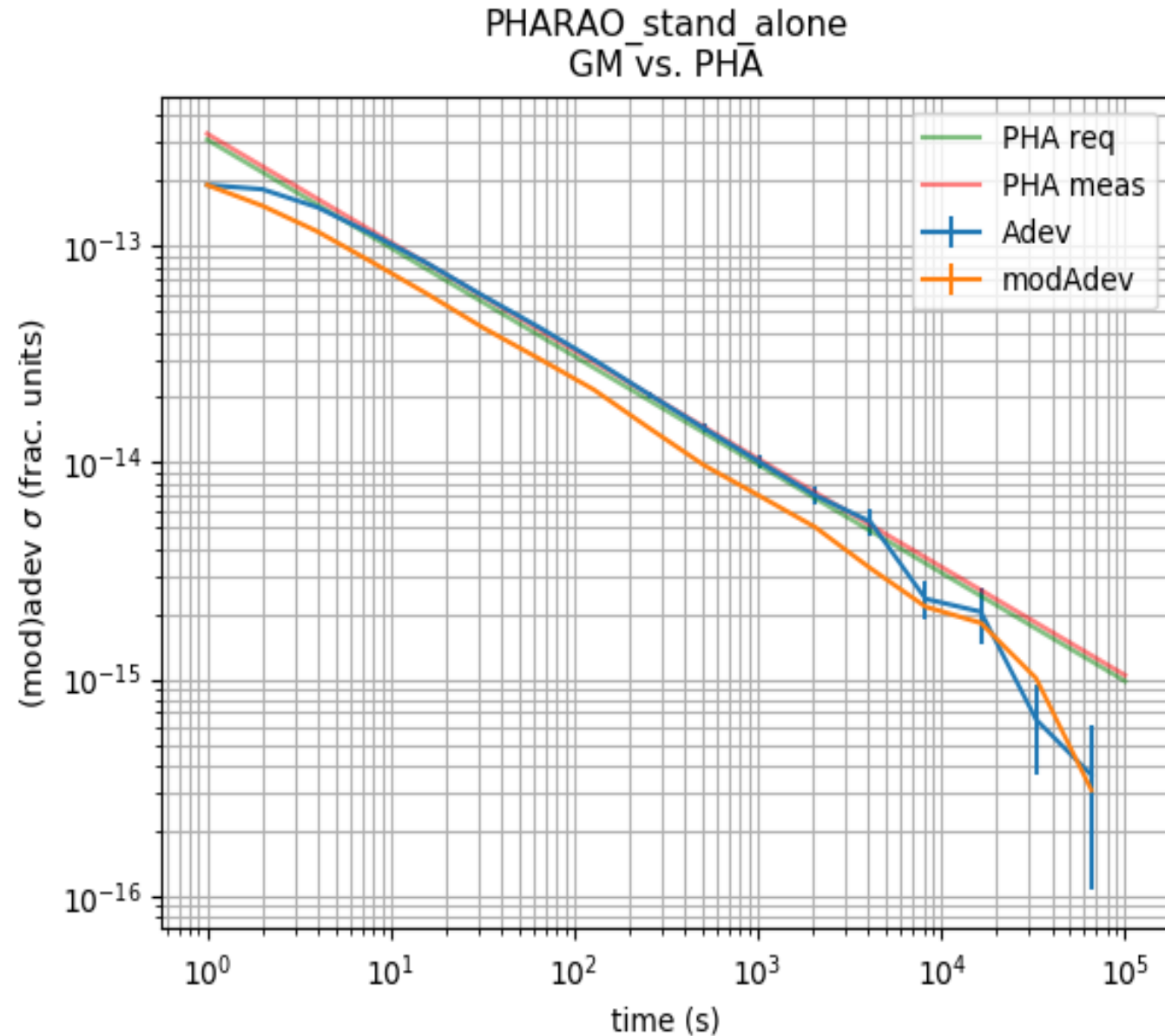
18.09.2020 and 19.09.2020 / SHM C-Field Scans / Before and after degaussing



Stand alone instruments

SHM

Three cornered hat measurement
SHM well within requirement
Degaussing works nicely.



PHARAO stability

As expected. Measurement
against external maser

Stand alone environment

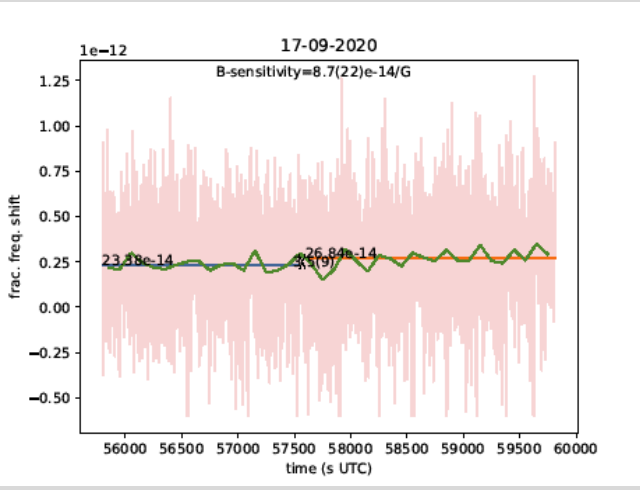
Subtitle

DEFENCE AND SPACE

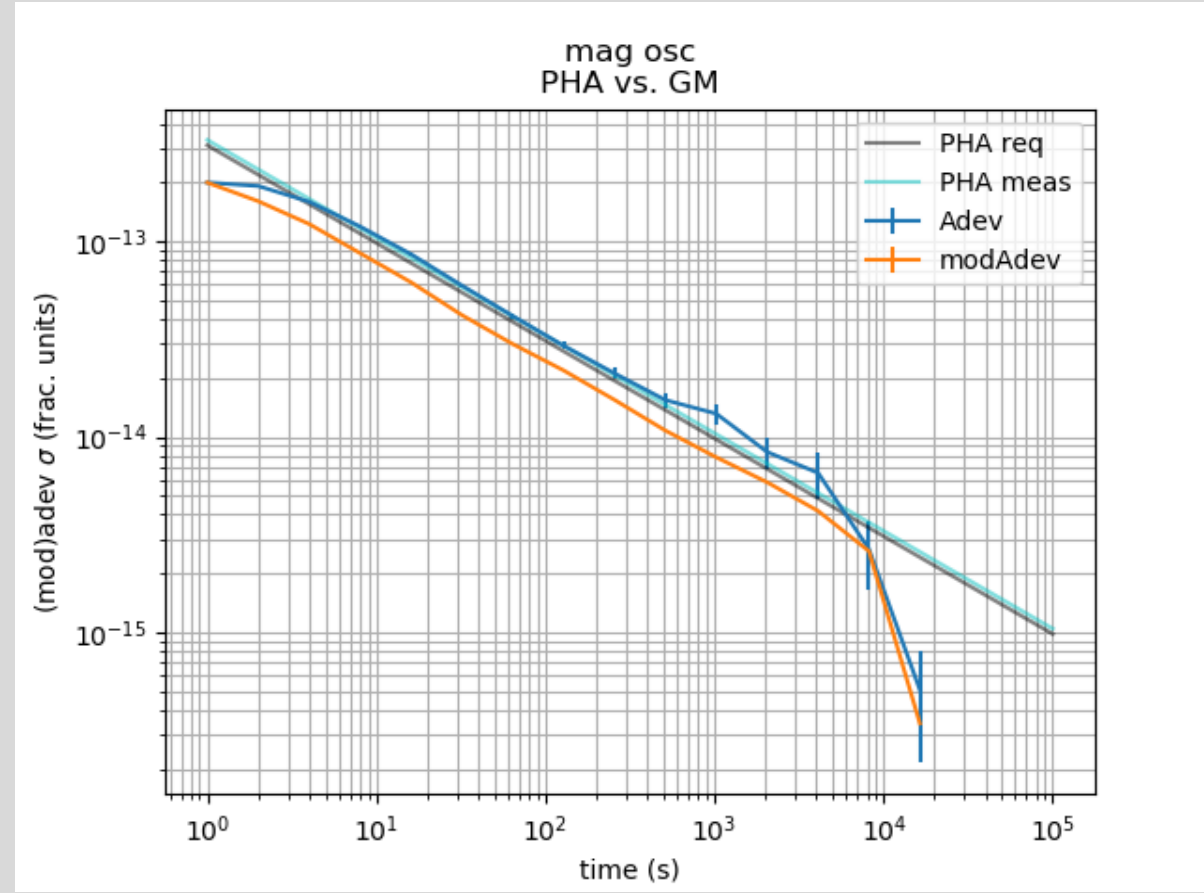
Stand alone environment

Magnetic PHARAO

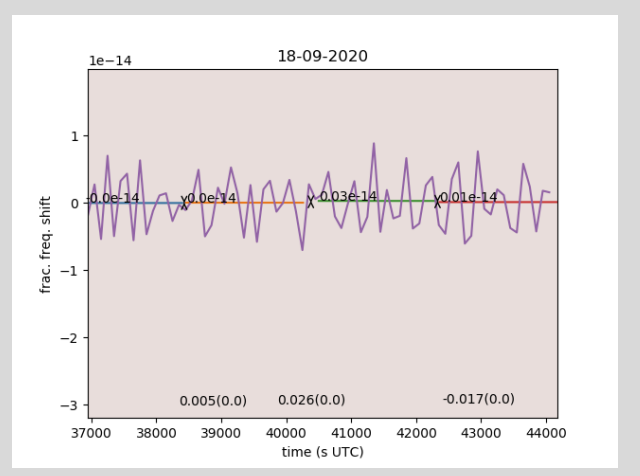
- PHARAO is barely susceptible (respecting its noise level)
- In orbit it may be directly visible



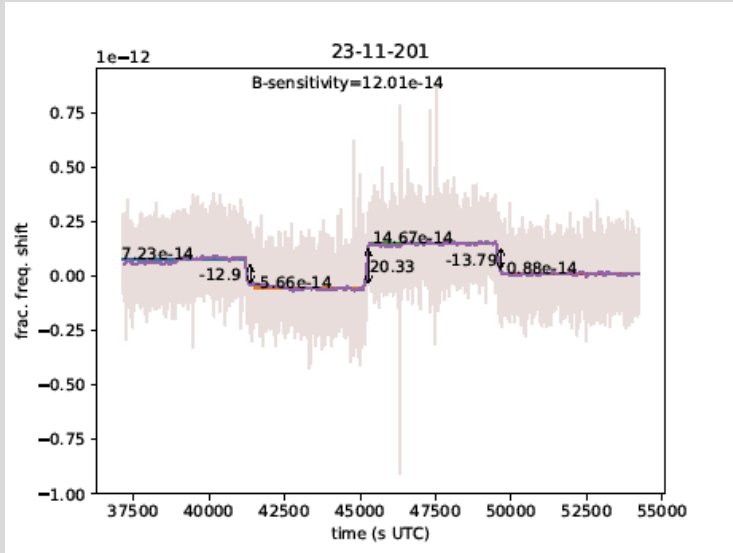
PHARAO axial
B sensitivity $8.7(2.2)e-14/G$



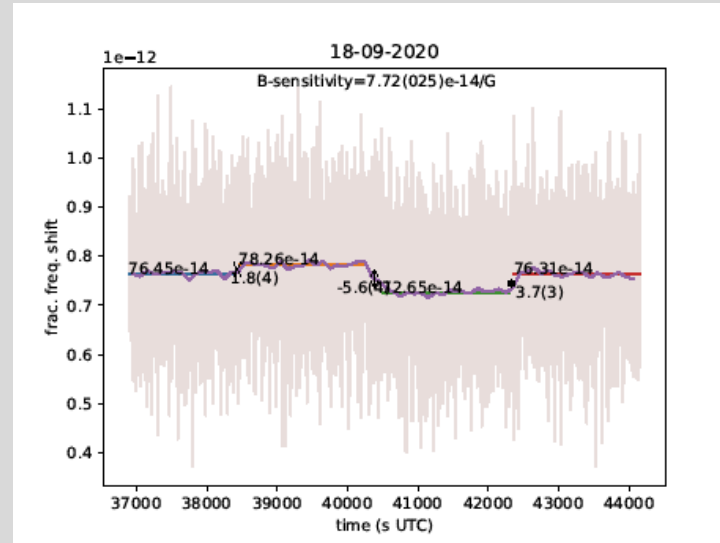
Magnetic orbital: 5000s



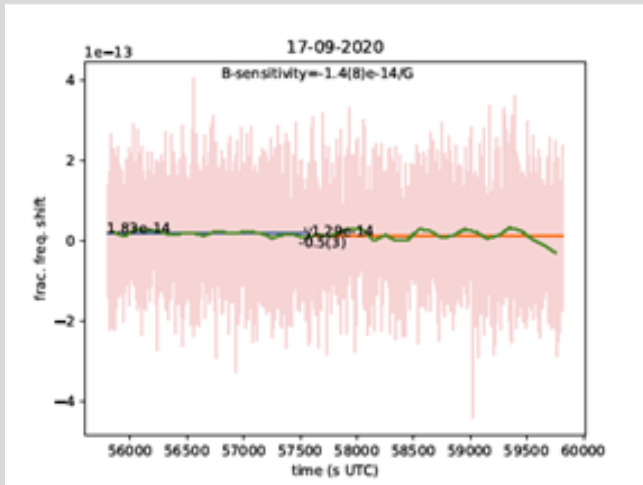
PHARAO radial
B sensitivity
(SHM measurement)



Stand alone reference: 1.2e-13/G at earth field pm 1 Gauss



ACES 7.73(25)e-14/G with 0 pm 0.35 Gauss



Radial B susceptibility 1.4(8)e-14

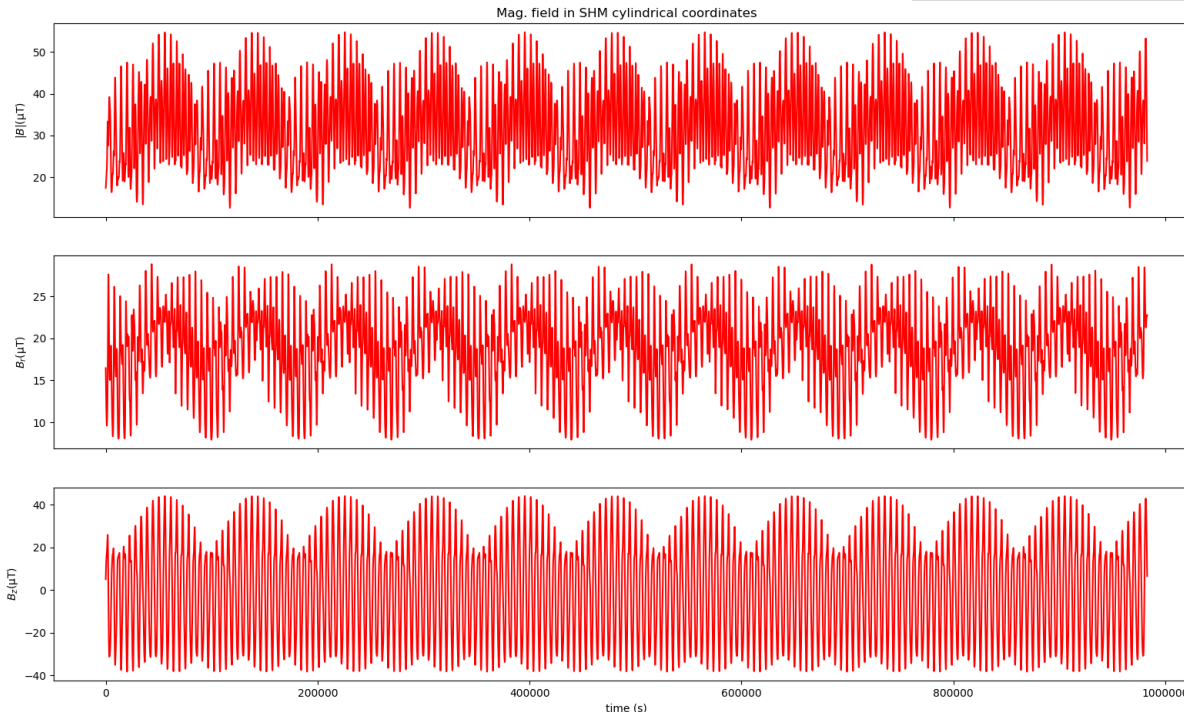
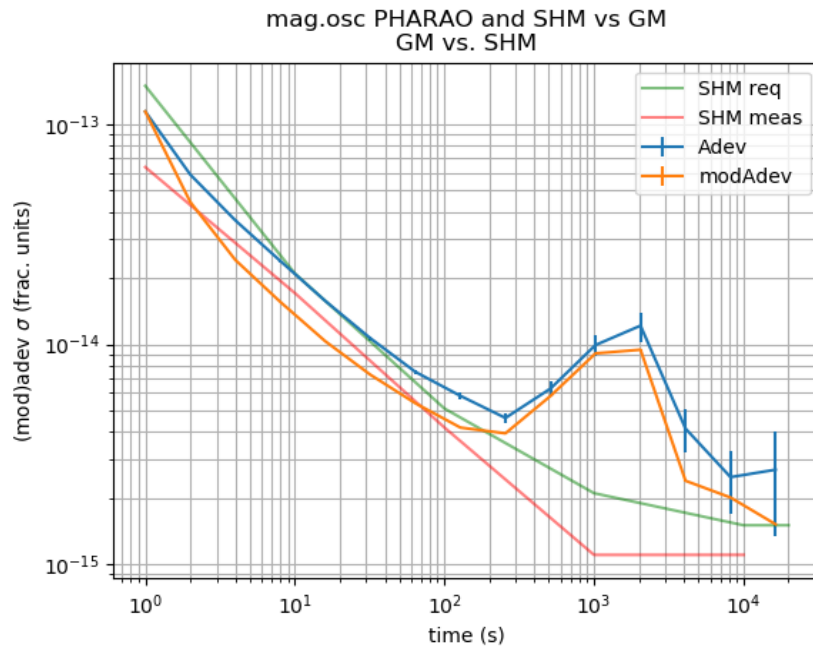
Stand alone environment

Magnetic SHM

- SHM after degaussing in zero field is less susceptible (shielding of vac chamber already included)
- SHM is clearly susceptible

Stand alone environment

Magnetic SHM

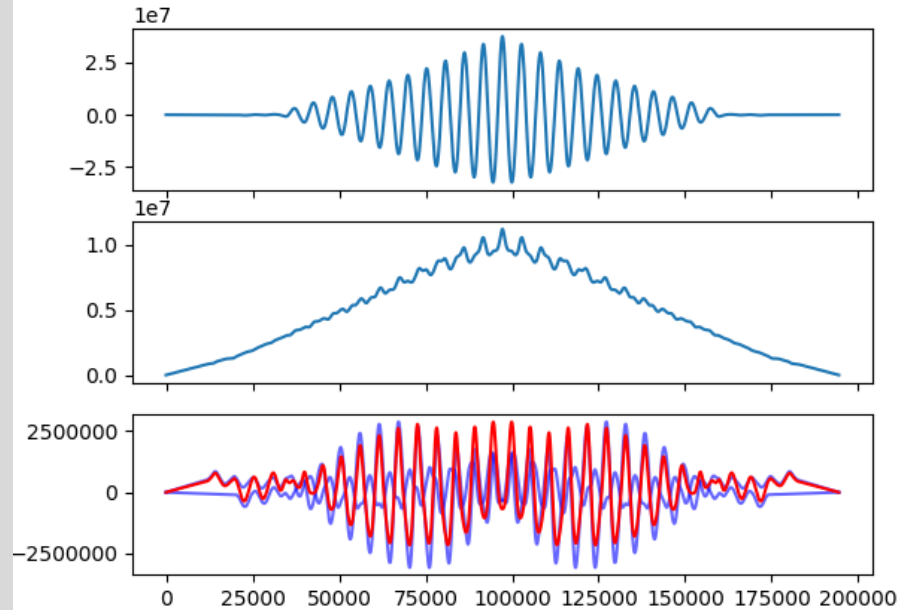
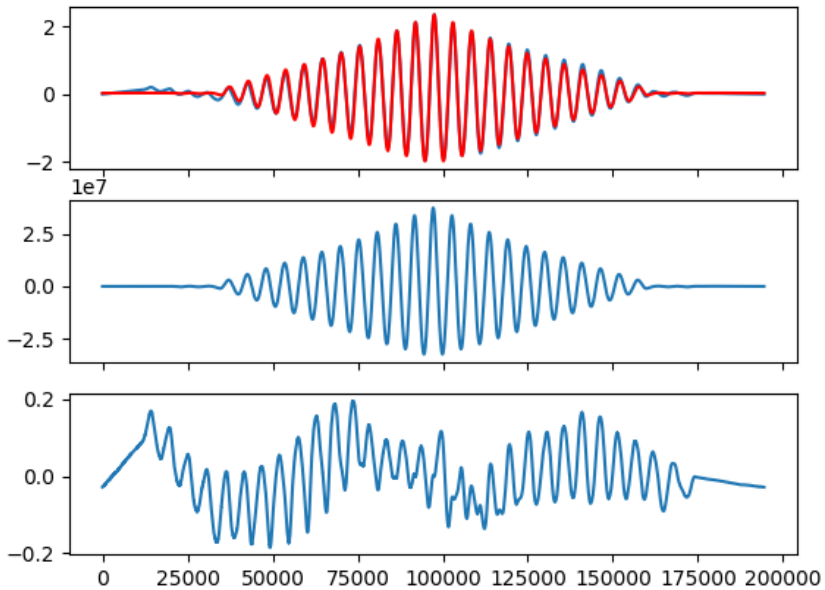


Total field

Radial to SHM

Axial to SHM

- Magnetic oscillation clearly seen in ADEV
- Magnetic field known in 3d and time



From top to down:

- Data with fit: $5.7640(15)e-14/G$
- Auto correlation
- residual

Same result as direct fit
Residual not strong
enough to fit radial
effect

From top to down:

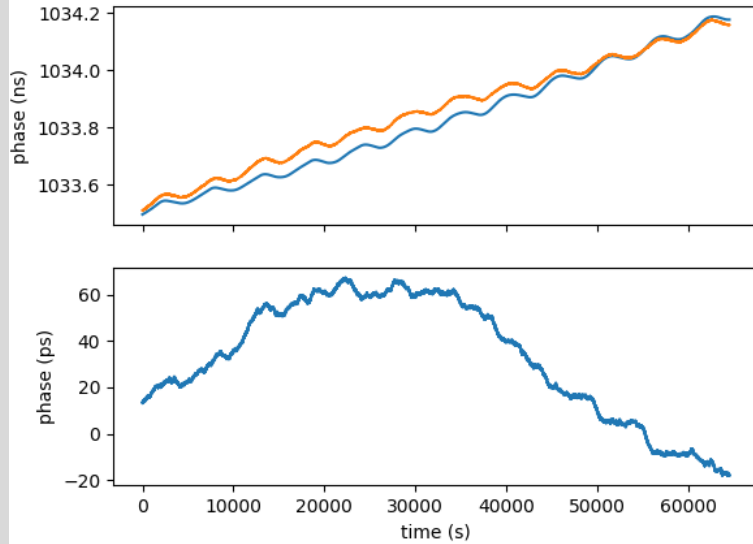
- Axial auto correlation
- Radial auto correlation
- Cross correlation axial and radial:
in red $B_z \cdot B_r + B_r \cdot B_z$

Stand alone environment

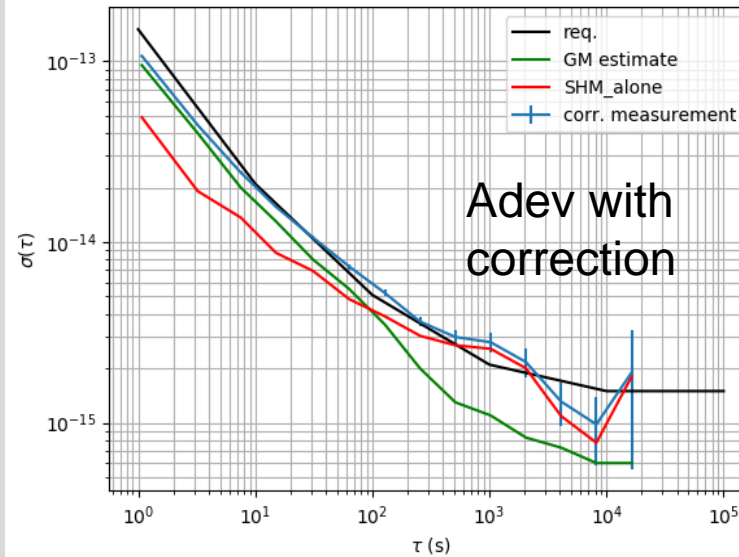
Magnetic SHM

Correction cross correlation

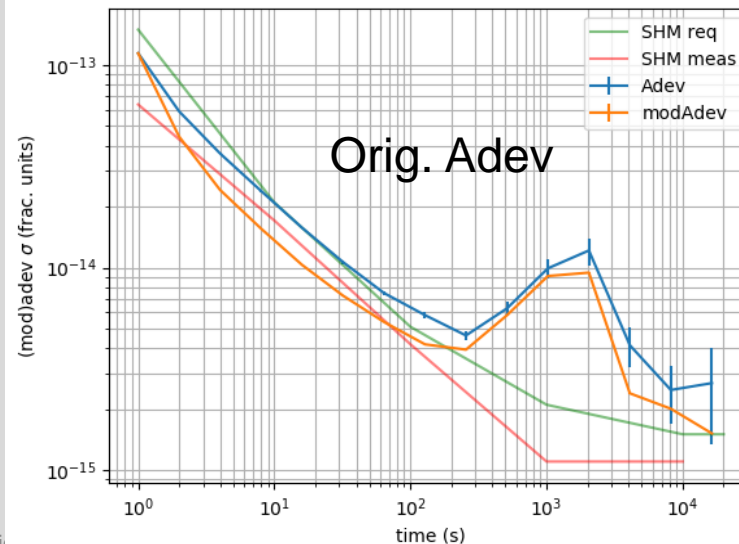
Same result, somewhat more precise. Radial effect visible



Integrate and apply frequency correction to raw PCO data:
 Top: data (blue) and correction (orange)
 Bottom: Phase track after removing drift and correction



mag.osc PHARAO and SHM vs GM
 GM vs. SHM



Stand alone environment

Magnetic SHM

Correction

- ADEV is nearly in requirement
- 5 fold suppression of the magnetic susceptibility
- B vector components must be known
- Not sensitive to angle of orbital simulator (pm 10° did not change final result)

Heater installed to compensate environmental conditions are used to induce expected residual effects on the instruments

Limited by heater power

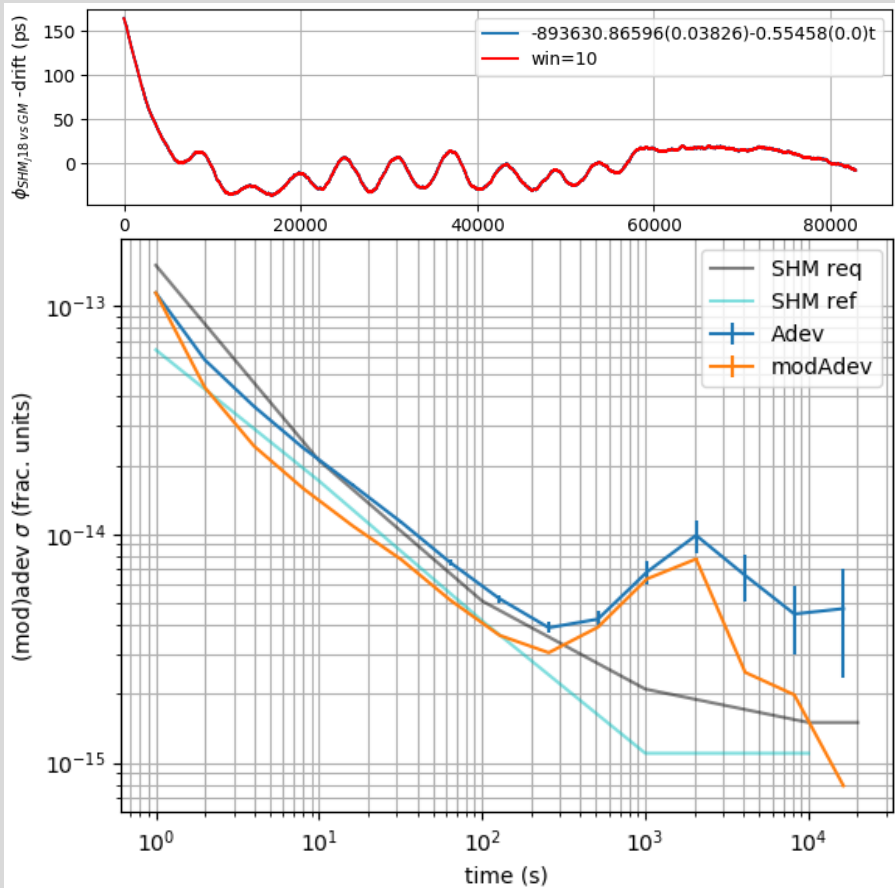
Instrument	Low temp required (°C)	Low temp reached (°C)	Amplitude required (pp°C)	Amplitude reached (pp°C)
PHARAO	14.9 (av TP)	$(17.5+21)/2=19.2$	3	2.5
SHM	21.2	24	1.5	1.5
FCDP	14.9	15	3.5	3.5

Limited by cooling power of shroud

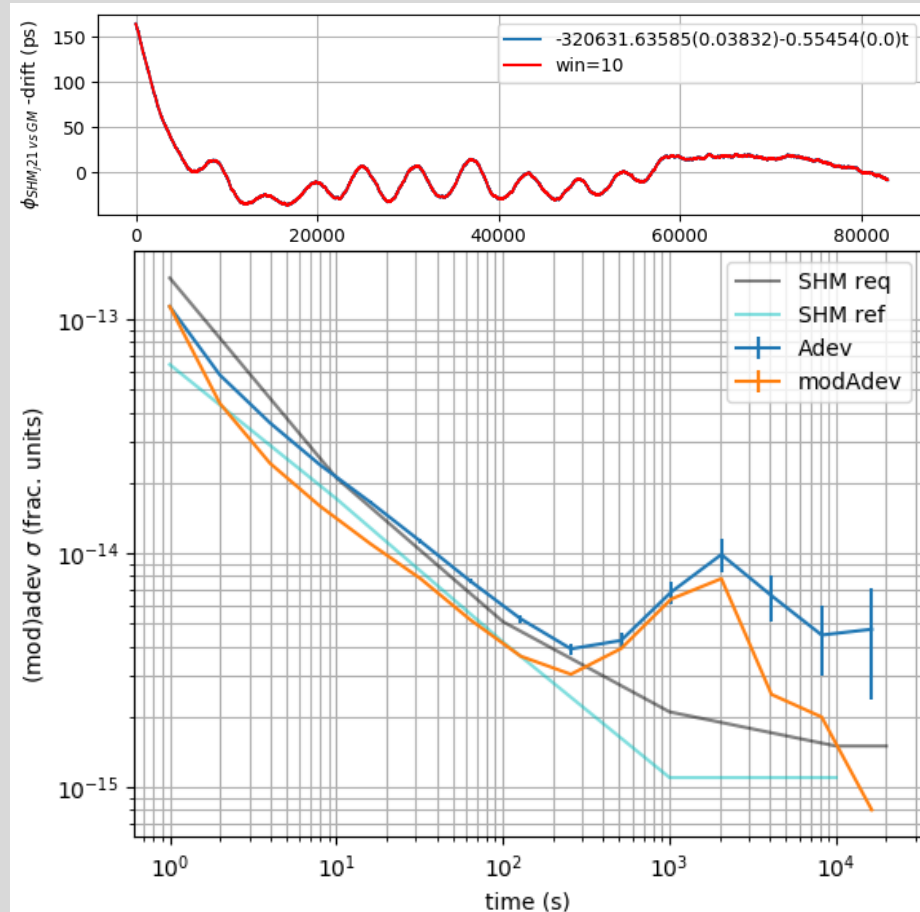
Stand alone environment

Temperature requirements

Improvements necessary
Difficult to implement



SHM through FCDP MWL output



SHM direct

Stand alone temperature

Thermal SHM

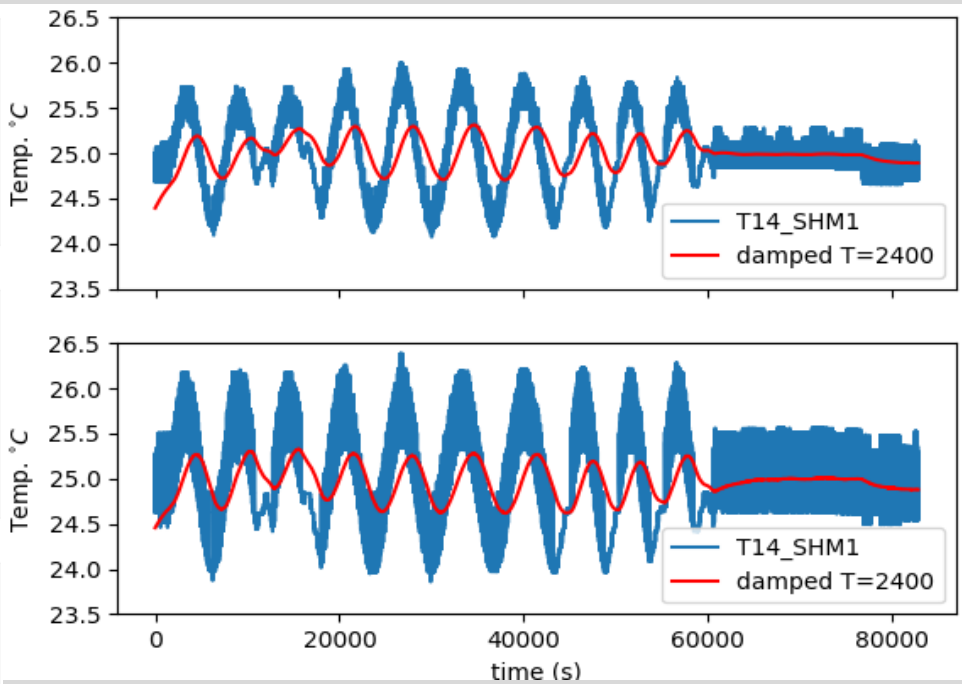
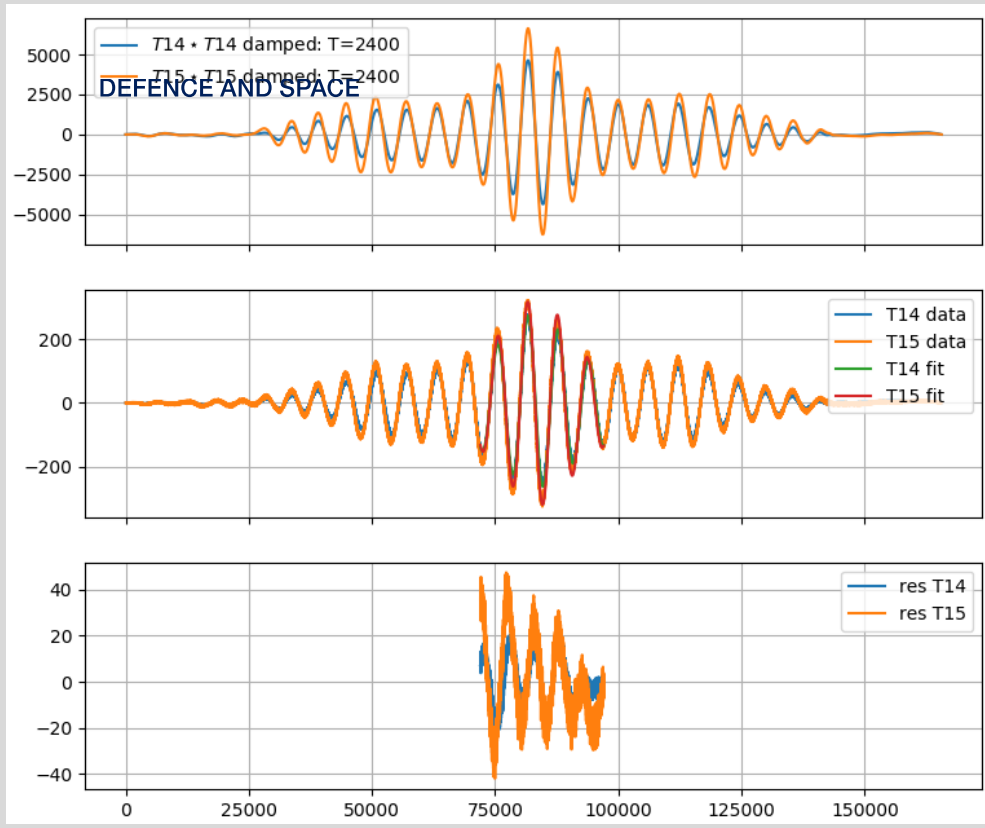
- Large thermal bump as expected
- FCDP has little impact on ADEV

Stand alone temperature

Thermal SHM

Cross correlation with damping

- Small residuals – good model
- Results roughly in agreement
- Fit error to small (residuals)

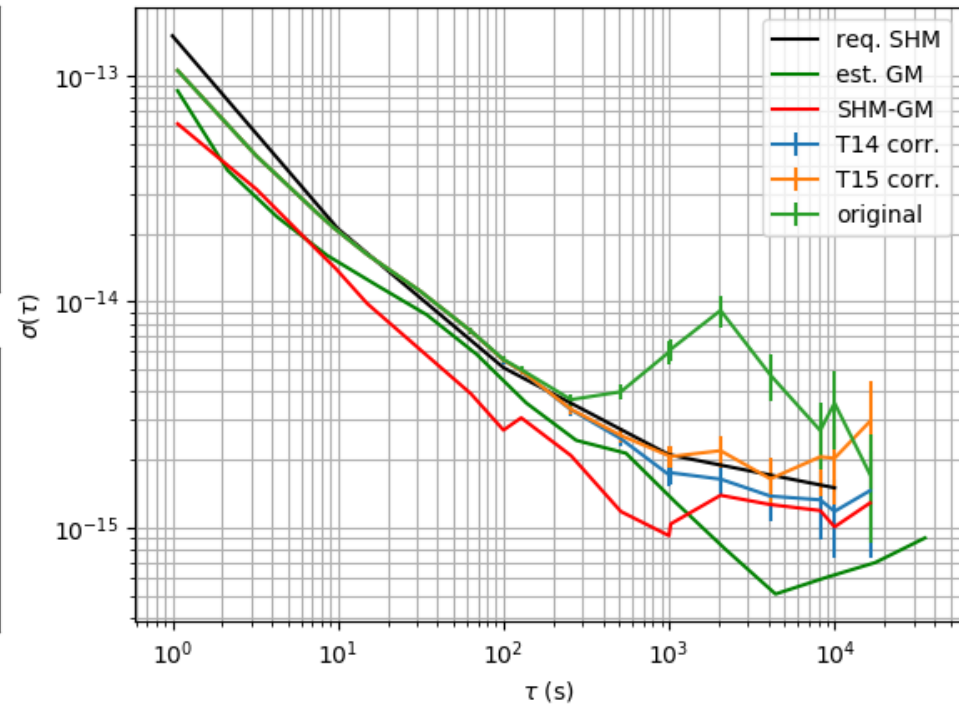
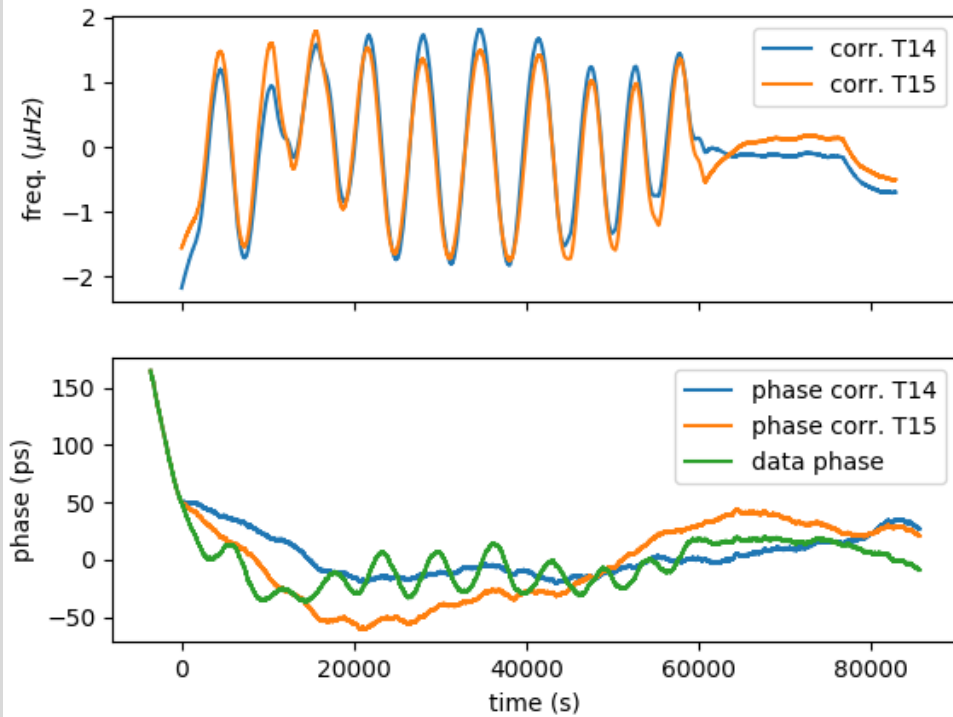


Damped temperature oscillations using convolution with:

$$\text{Simple RC damping: } \frac{X}{T} e^{-\frac{t}{T}}$$

Cross correlation with damped temperature and frequency data

Data	Amplitude (Hz/K s)	Damping T (s)	Offset (μHz/K s)	χ_T (μHz/K)
T14_SHM1@J21	2.481(1)e-05	2.431(3)e+03	-2.64(6)e+00	6.031(7)
T15_SHM2@J21	1.994(2)e-05	2.799(3)e+03	-2.9(1)e+00	5.581(8)
T14_SHM1@J18	2.471(1)e-05	2.45000000(8)e+03	-2.58(7)e+00	6.054(2)
T15_SHM2@J18	1.984(2)e-05	2.822(4)e+03	-2.8(1)e+00	5.60(1)



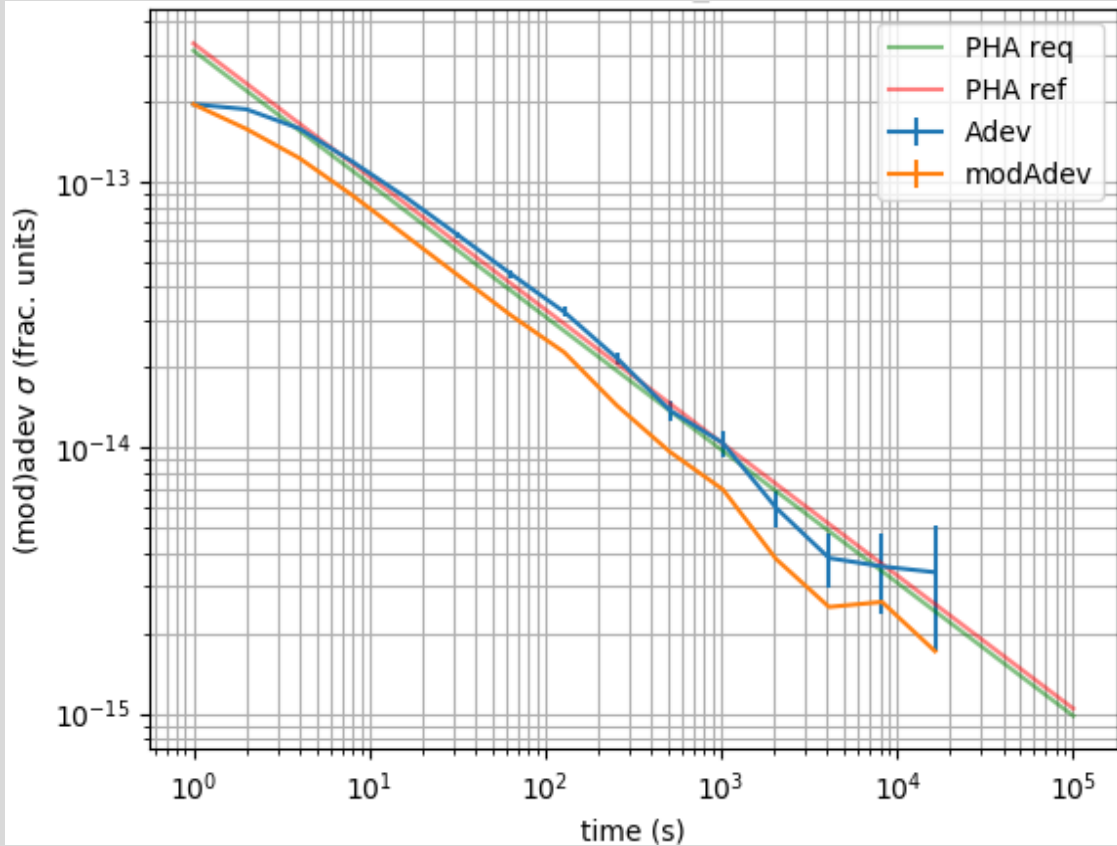
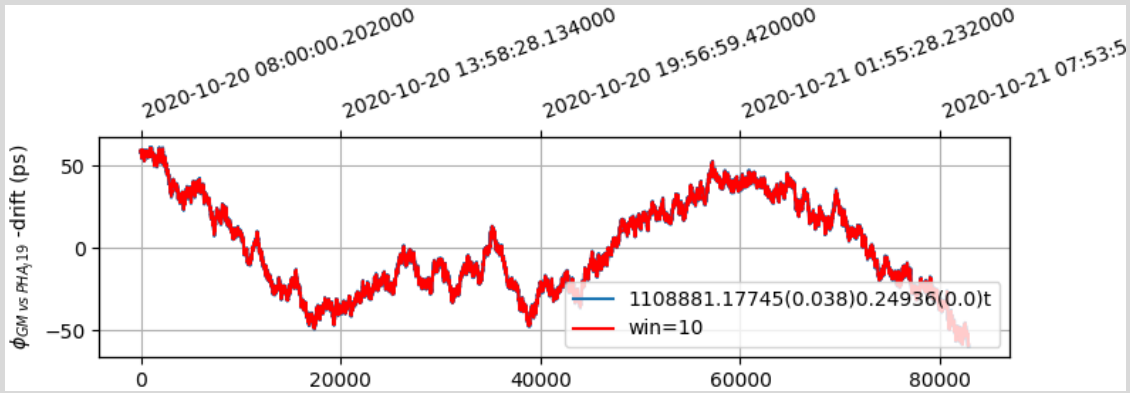
Stand alone temperature

Thermal SHM

Checking results on data

χ_T ($\mu\text{Hz}/\text{K}$)	Damping T (s)
6.031(7)	2.431(3)e+03

- SHM can be corrected to be within requirement
- Influence of temperature sufficiently understood for calibration
- T14 gives a better estimate

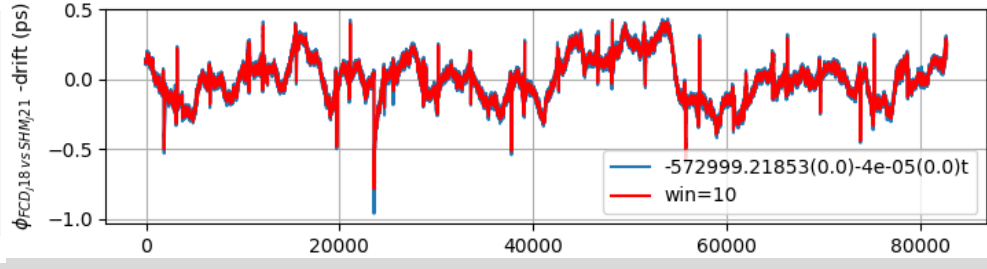
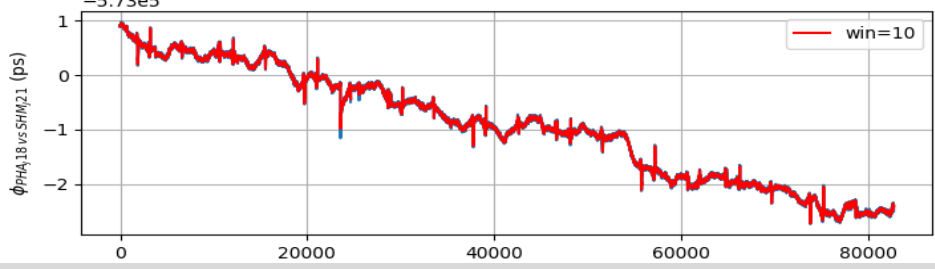


Stand alone temperature

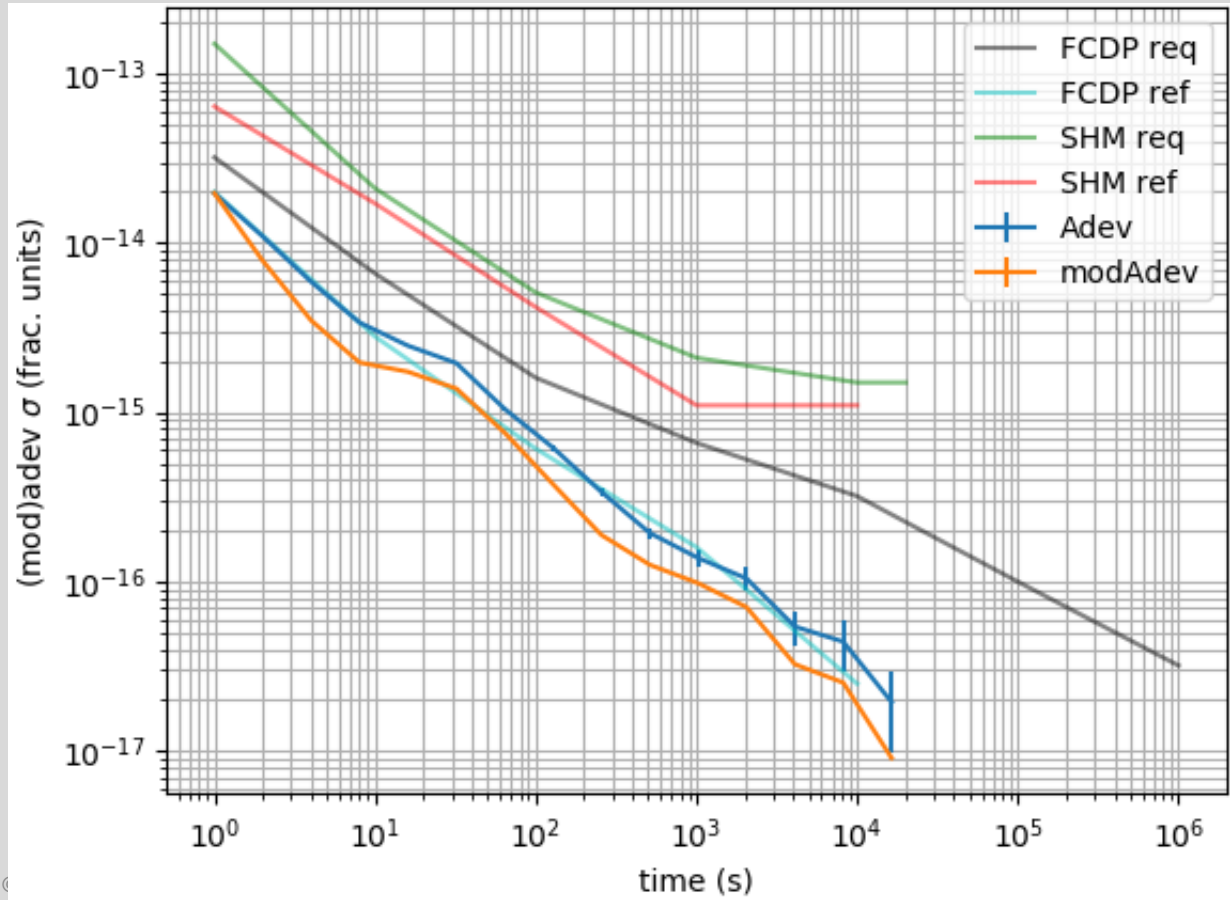
Thermal PHARAO

- PHARAO is not affected by temperature variations for stand alone

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SHM MWL vs SHM, FCDP by pass

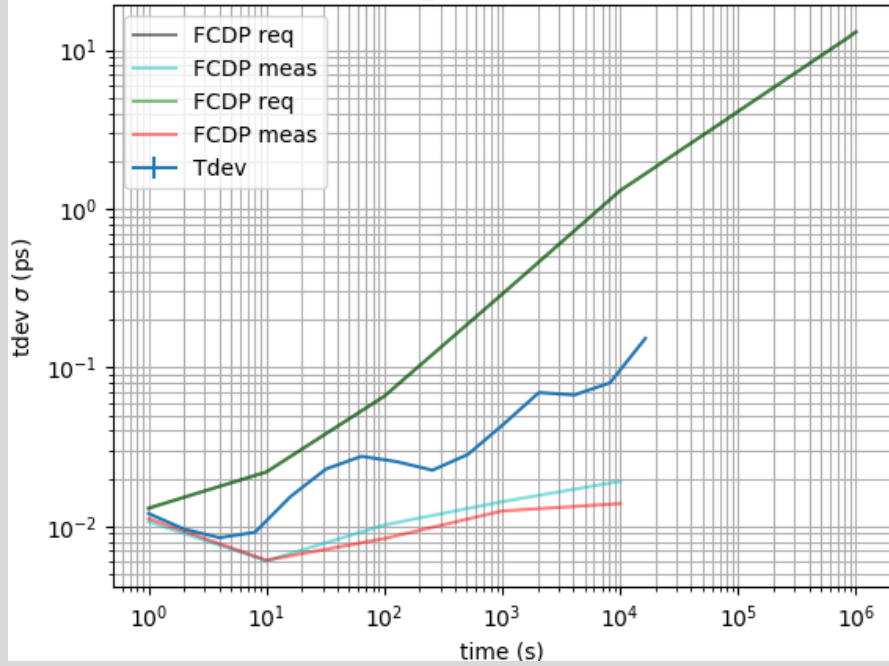


Stand alone temperature

Thermal FCDP

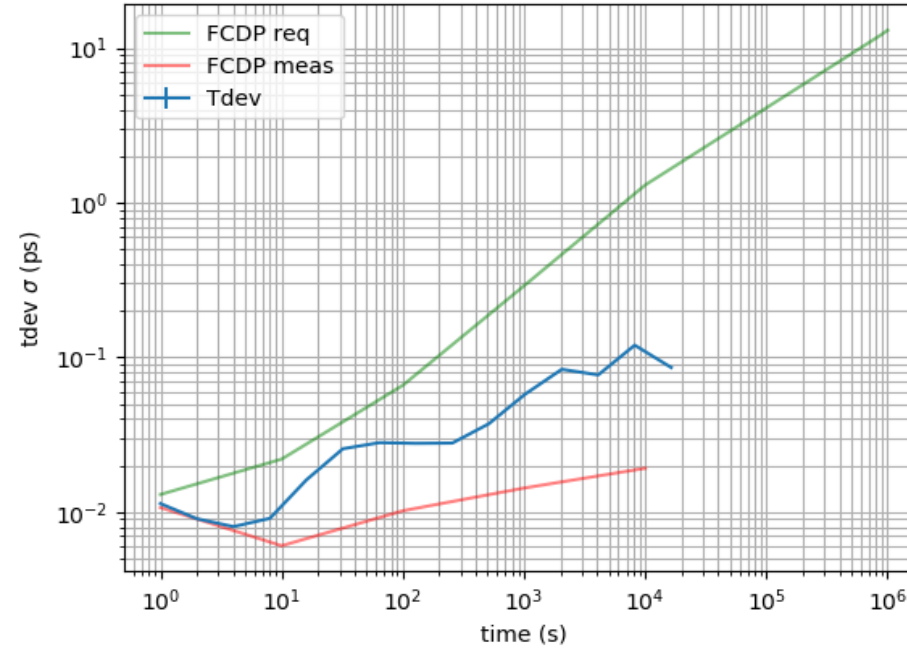
- ADEV is not impacted

stand alone: Thermal oscillations
SHM J18 vs. SHM J17



MWL vs Test output

stand alone: Thermal oscillations
SHM J21 vs. SHM J18



MWL vs FCDP bypass

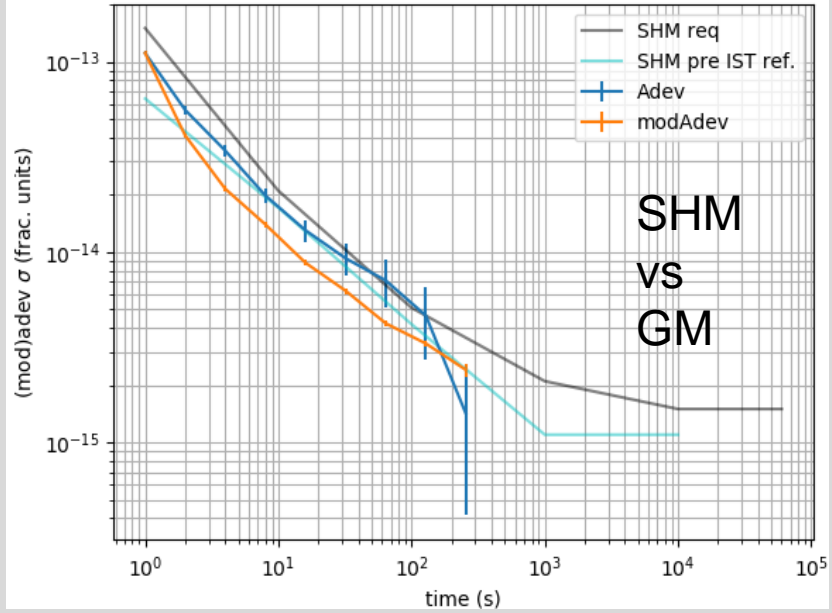
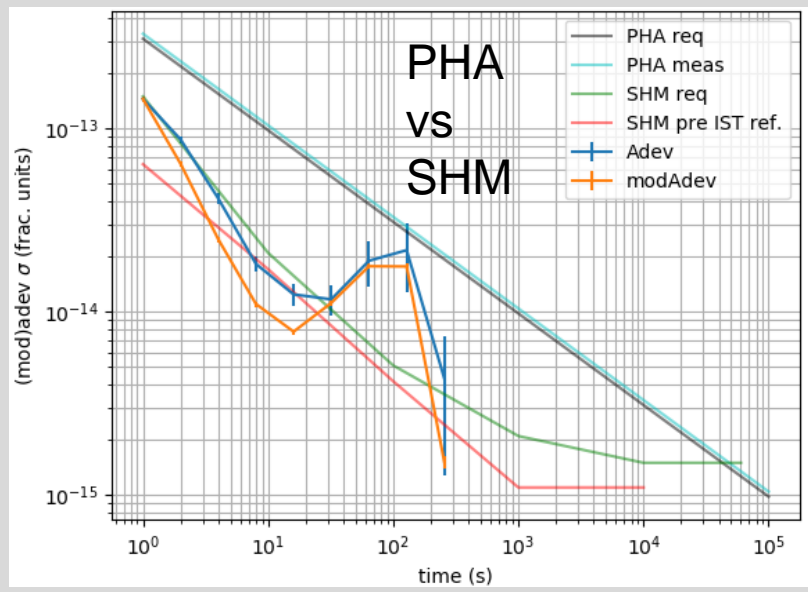
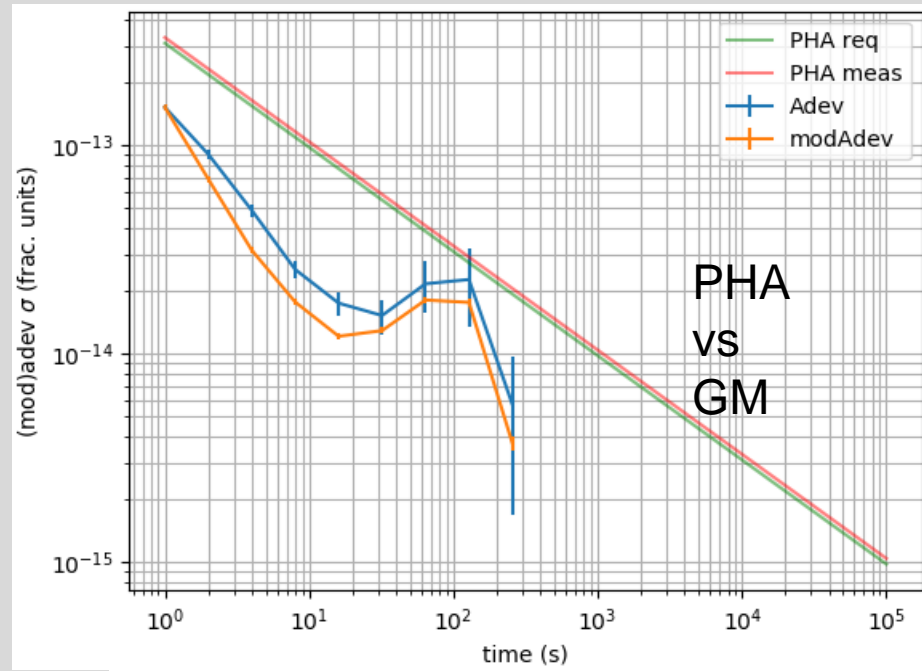
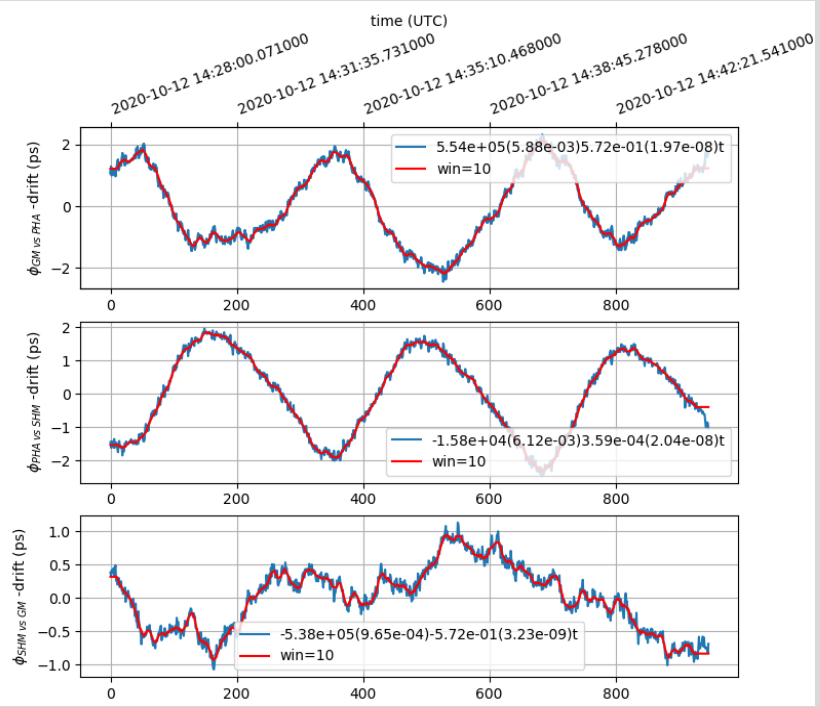
Stand alone temperature

Thermal FCDP

TDEV in requirement

Short Term Servo Loop

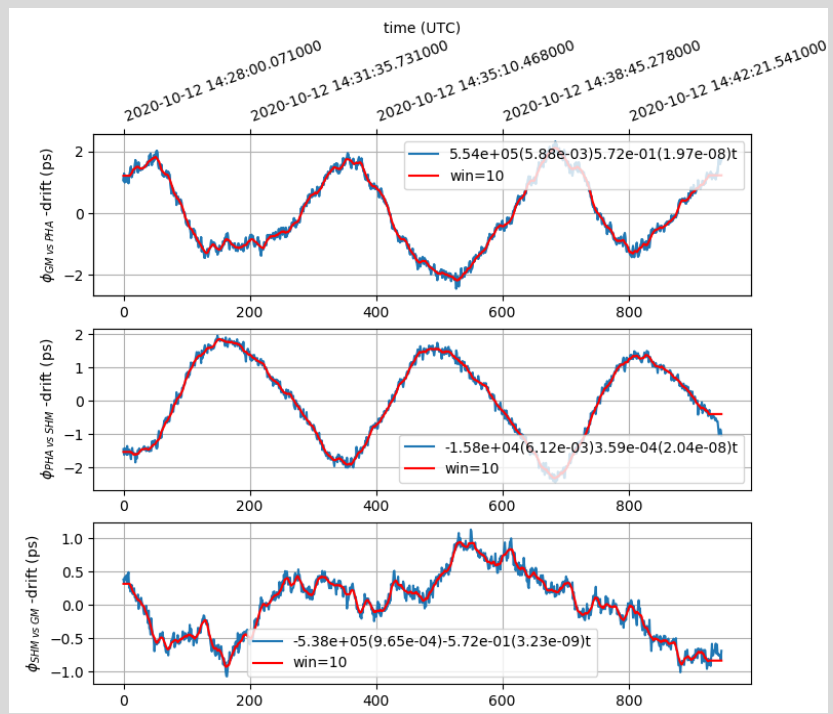
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STSL

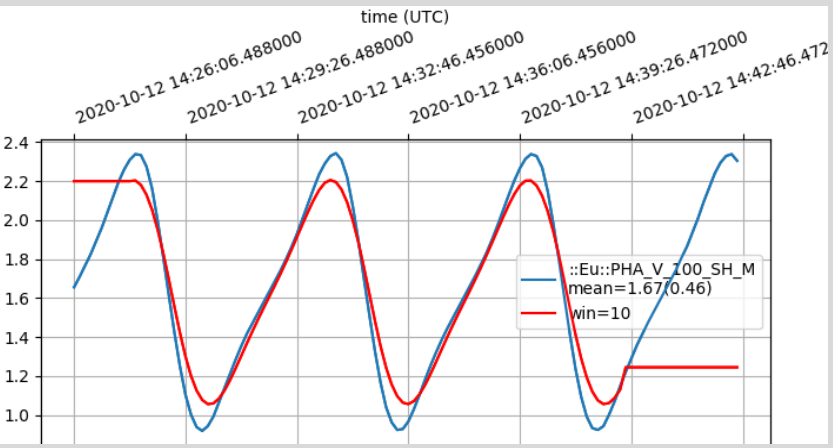
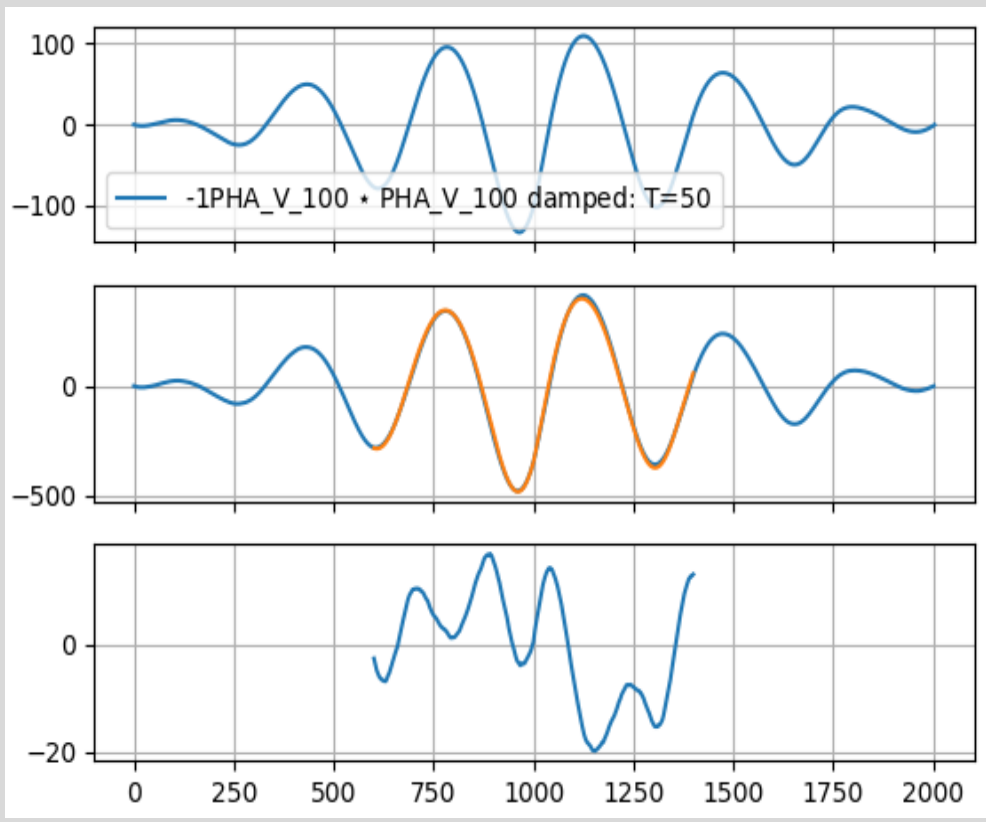
First Data AMPM

- STSL works in principle
- AMPM is a clear problem



Cross correlation with damped effect of the voltage:

$$\frac{X}{T} e^{-\frac{t}{T}}$$

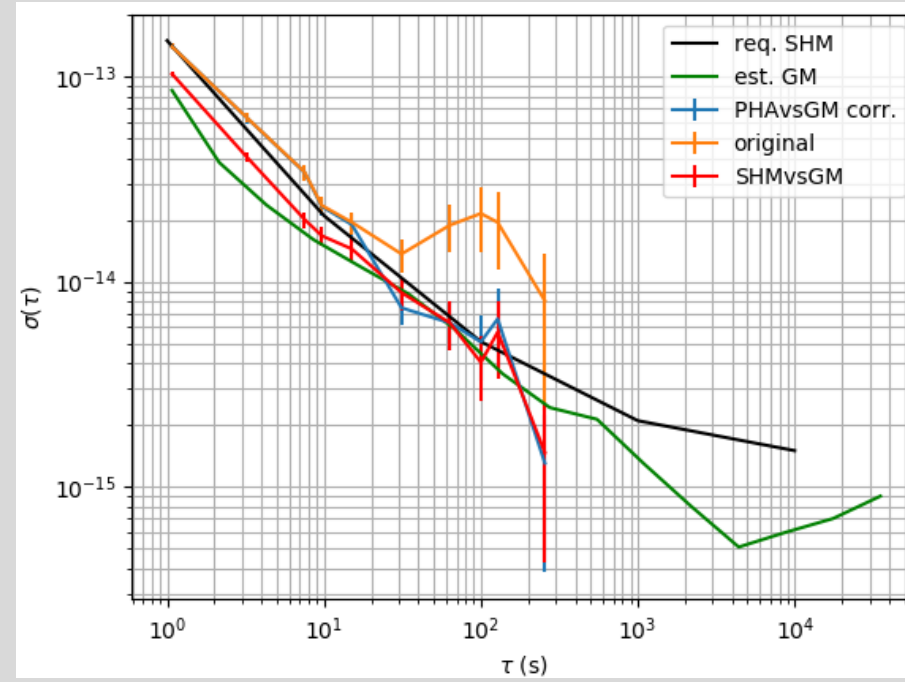
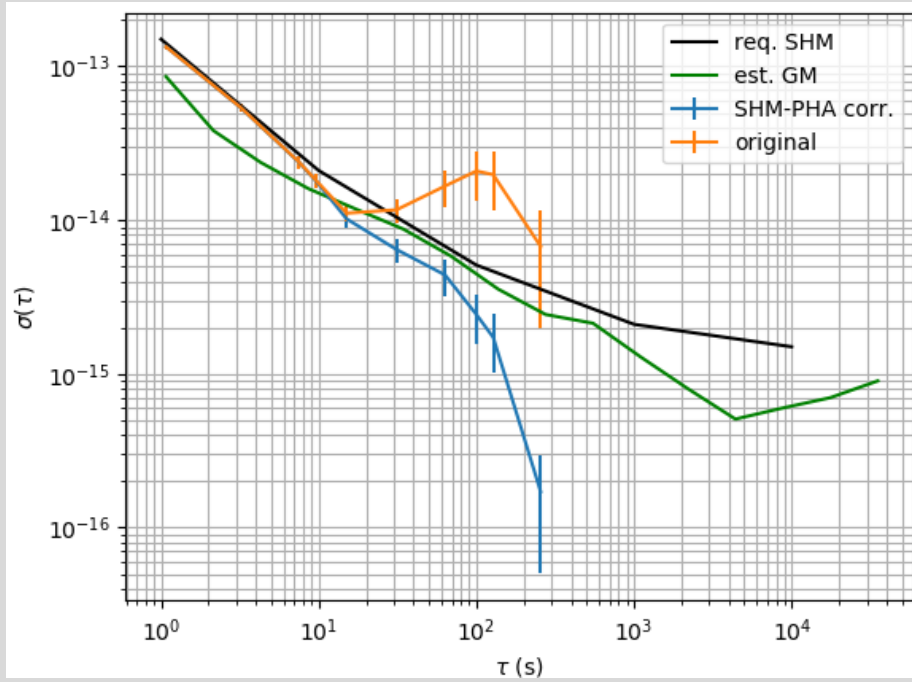


STSL

STSL correction

- Cross correlation shows good match
- Residual is no more than 4% of signal
- Residual appears to oscillate at a higher frequency

Rapid oscillations due to temperature sensitivity of PLL and oscillating temperature controller (HW)



STSL

Correction with
average parameter:
Calibration

- Calibration is sufficient

Magnetic orbital simulation with Short Term Servo Loop

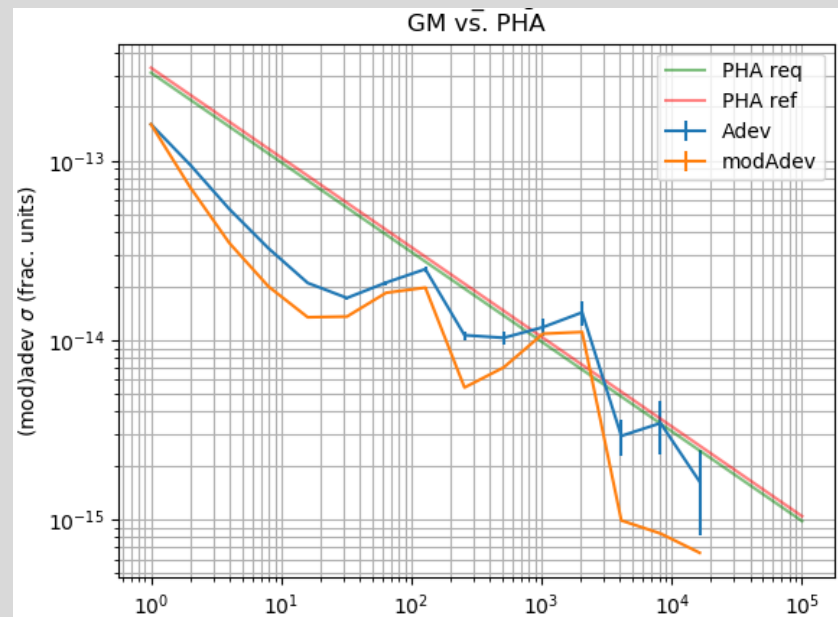
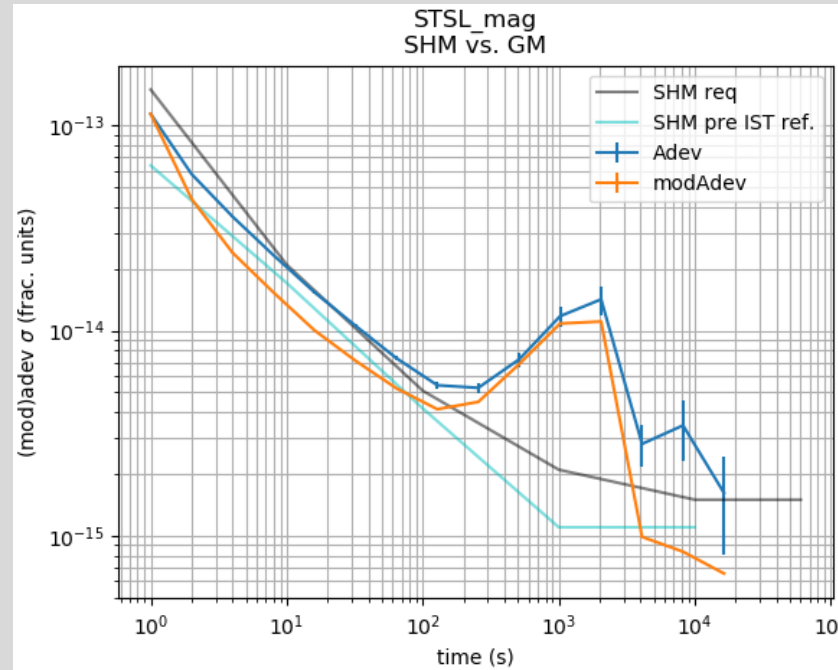
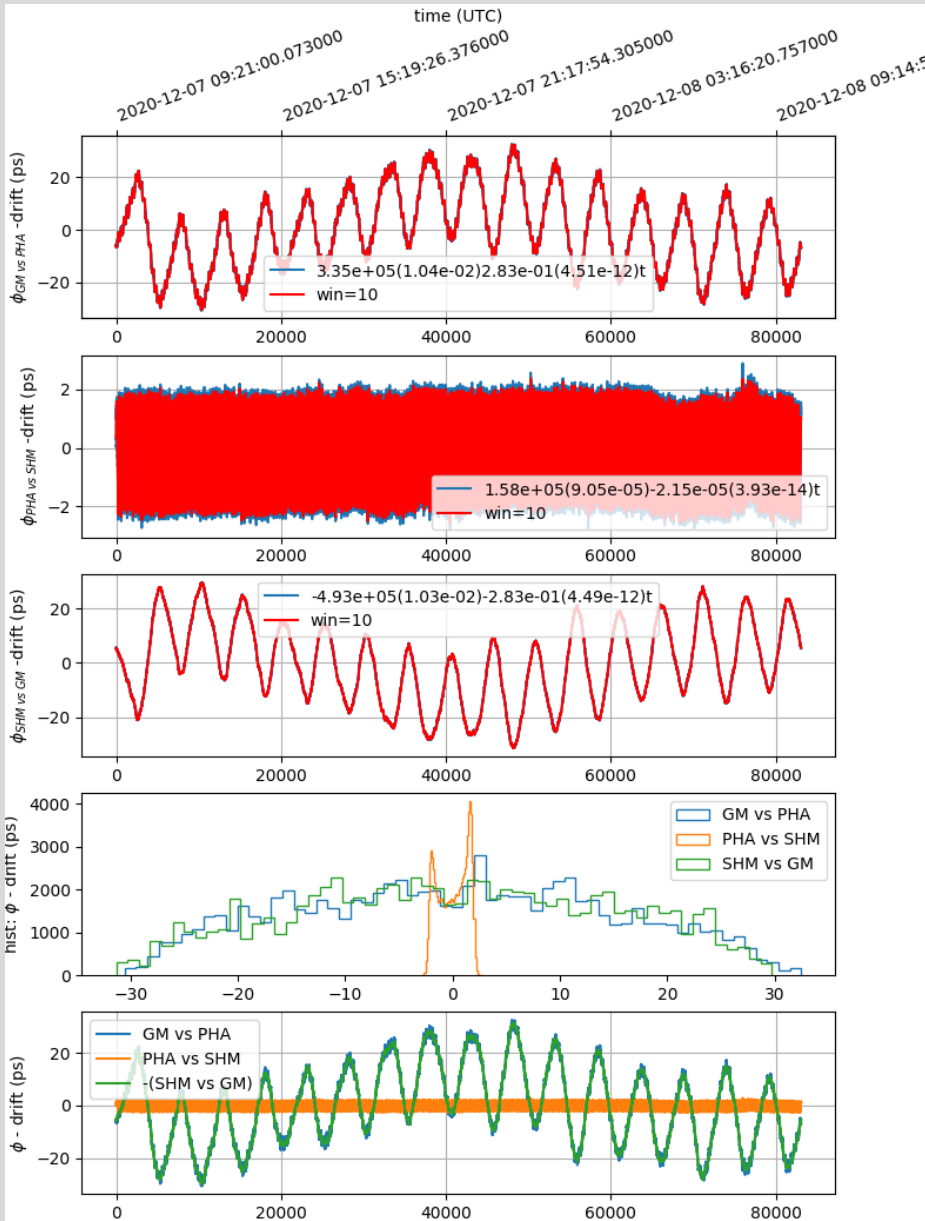
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AIRBUS

STSL magnetic orbit

Data

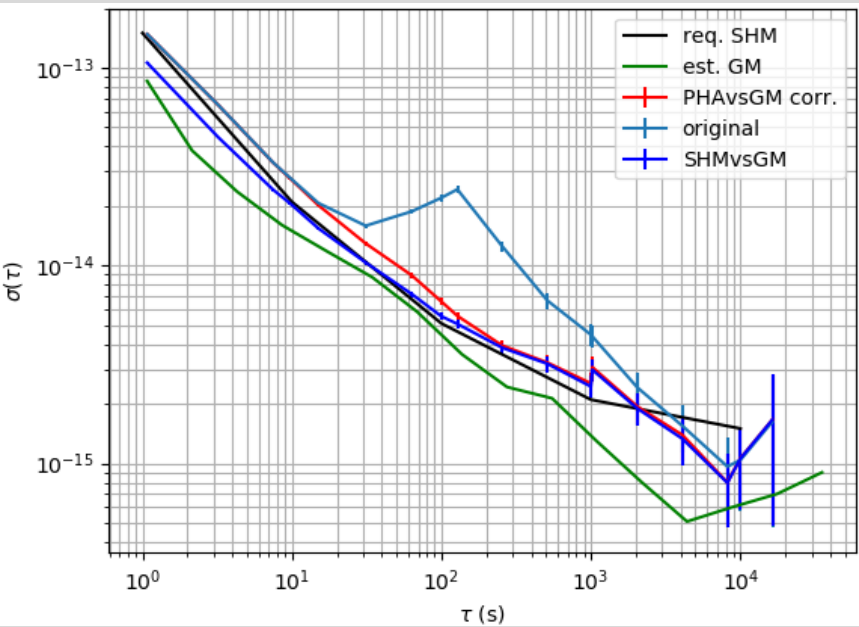
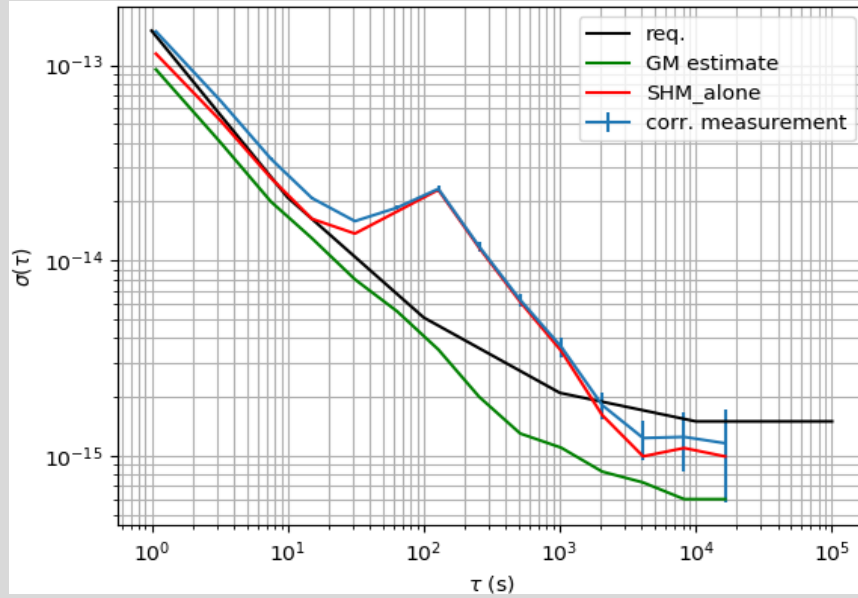
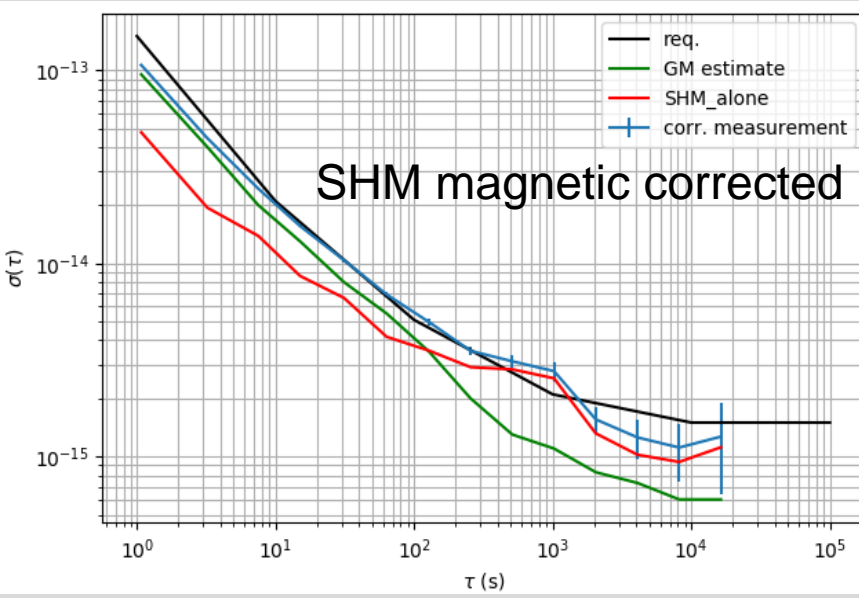
- Large magnetic bump from SHM
- AMPM conversion on top of PHARO
- No influence on AMPM conversion



STSL magnetic orbit

Corrections applied

- 5.7720(4)e-14/G correction (within 1% of previous measurement)
- AMPM conversion factor near previous point
- Both corrections successful
- Slight increase in final ADEV



Long Term Servo Loop

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25 Sets of LTSL parameters tested

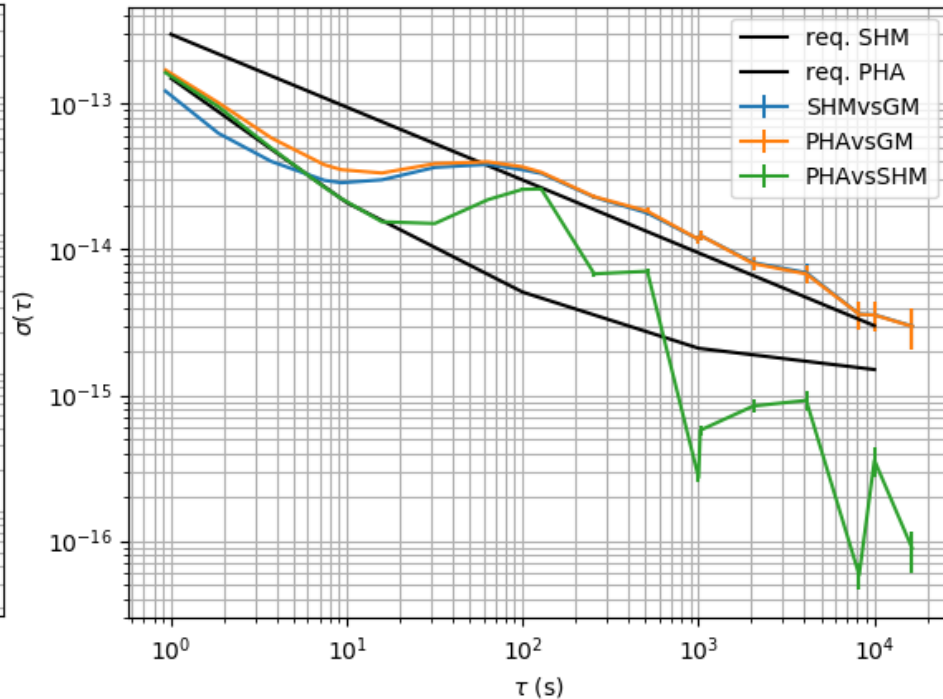
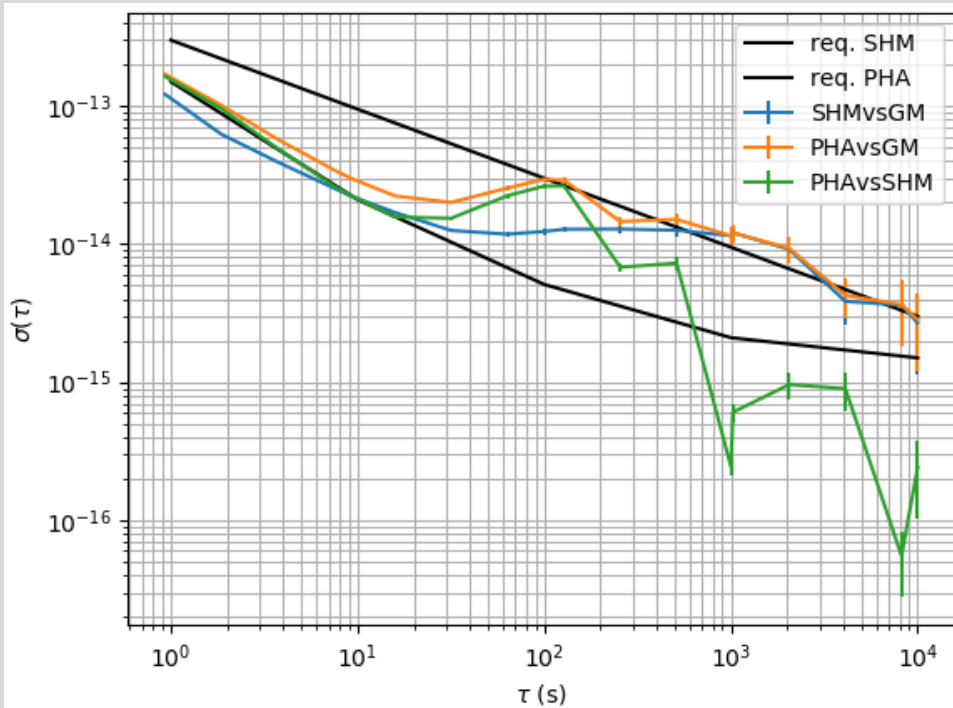
KA should be 0

KB small and negative (-0.0005)

KC small and positive (0.005)

KB=-0.0005 KC=+0.005

KB=-0.0025 KC=+0.01



LTSL

Full orbital simulation with STSL and LTSL

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AIRBUS

8 x 10 oscillations with 2 hours break in between
 Discriminating them from magnetic oscillations

Full orbital simulation

Temperature range

Cooler limitation

Instrument	High temp required (°C)	High temp reached (°C)	Amplitude required (pp°C)	Amplitude reached (pp°C)
PHARAO	29.5 (av TP)	$(24.85+26.4)/2=25.65$	3	2.5
SHM	33.4	30.2	1.5	1.5
FCDP	42.6	22.8	3.5	3.5
Instrument	High temp required (°C)	High temp reached (°C)	Amplitude required (pp°C)	Amplitude reached (pp°C)

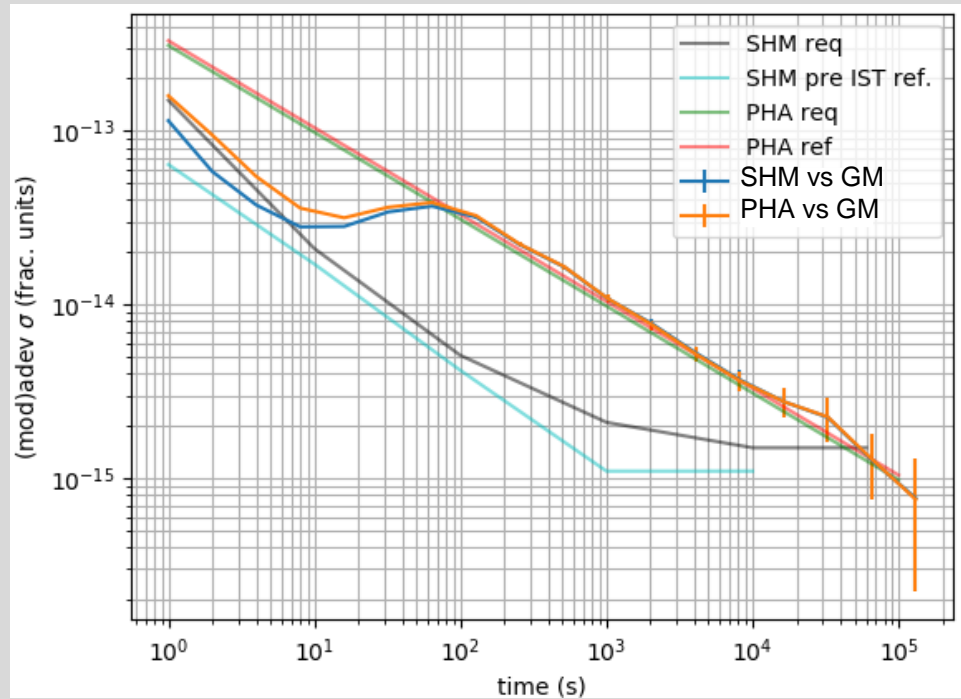
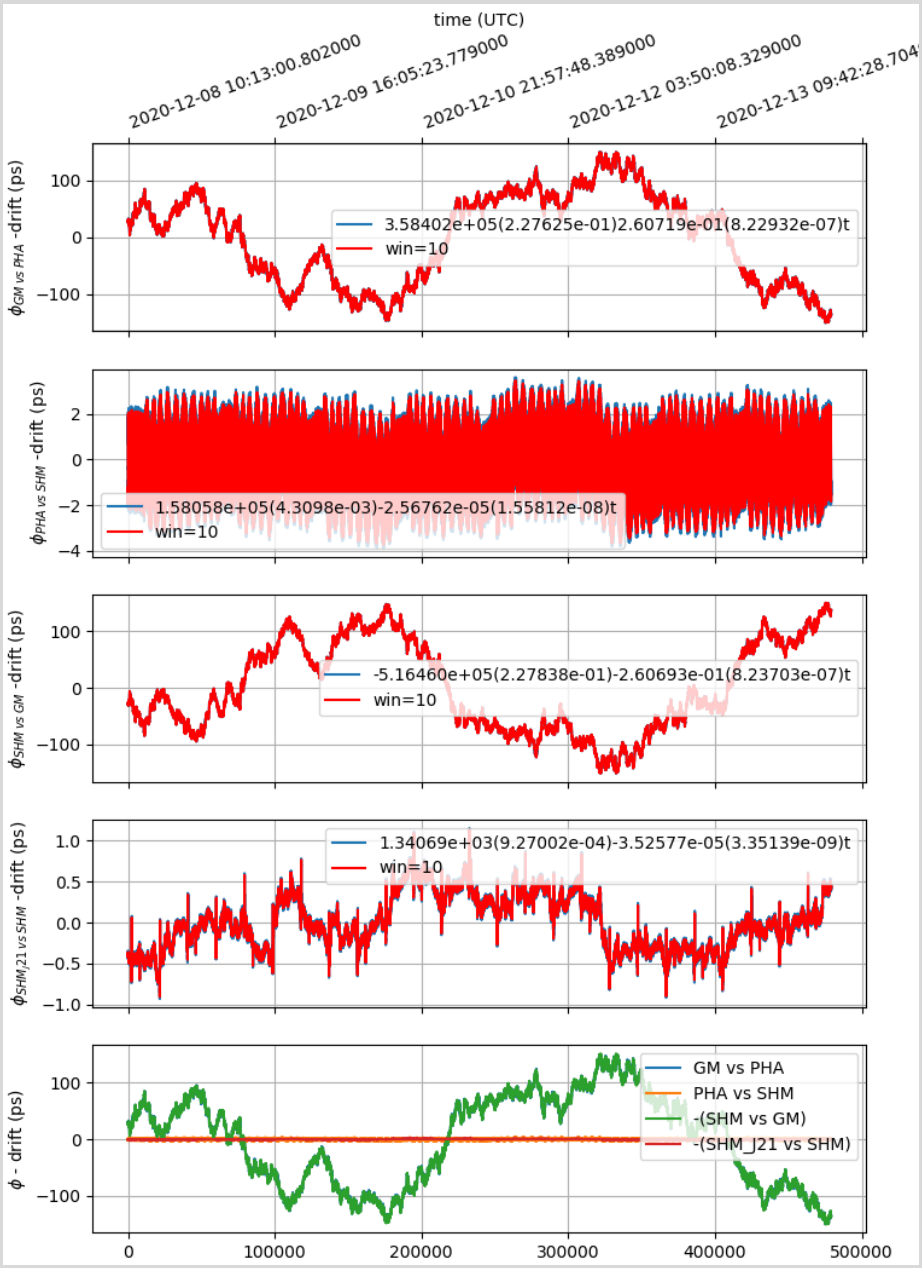
Heater/cooler limitation

Safety margin thermo stat

Full orbital oscillation

PCO data

- LTSL is covering AMPM
- Influence of orbital conditions not observed in ADEV



Nearly 6 days measurement
 STSL: Kp=40000 Ki=500
 LTSL: KA=0 KB=-0.00250 KC=+0.001

Quick Summary

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AIRBUS

Magnetic field:

SHM susceptibility is $5.76(5)e-14/G$ in axial direction.

PHARAO has no significant magnetic susceptibility (compared to on earth instability)

The orbital magnetic field can be matched to the SHM frequency difference to better than one minute using correlations and fitting.

Thermal variation

SHM susceptibility is $6.03e-13/K$ damping time 2430 s

PHARAO susceptibility insignificant (compared to on earth instability)

Setup needs to be improved (on the way)

LTSL

Best parameter set is $KA=0, KB=KC=0.0005$

STSL

Optimal parameter near $Kp=40000(-10000+5000)$ $Ki=500(-100+500)$

AMPM susceptibility $-3.65(11)$ ps/V damping time 48.2(3.7)s

AMPM susceptibility is temperature dependent (not shown 0.53 ps/K) and can be calibrated in orbit

SHM needed refurbishment due to vacuum problems

Setup needs improvement for temperature range

MWL will be included in next test

Lessons learned

Thank you

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